

# Solving high-dimensional parameter inference: marginal posterior densities & Moment Networks

Niall Jeffrey<sup>1,2</sup>, Benjamin D. Wandelt<sup>3,4</sup>

<sup>1</sup> *École Normale Supérieure, Paris*

<sup>2</sup> *University College London*

<sup>3</sup> *Institut d'Astrophysique de Paris (IAP), Sorbonne Université*

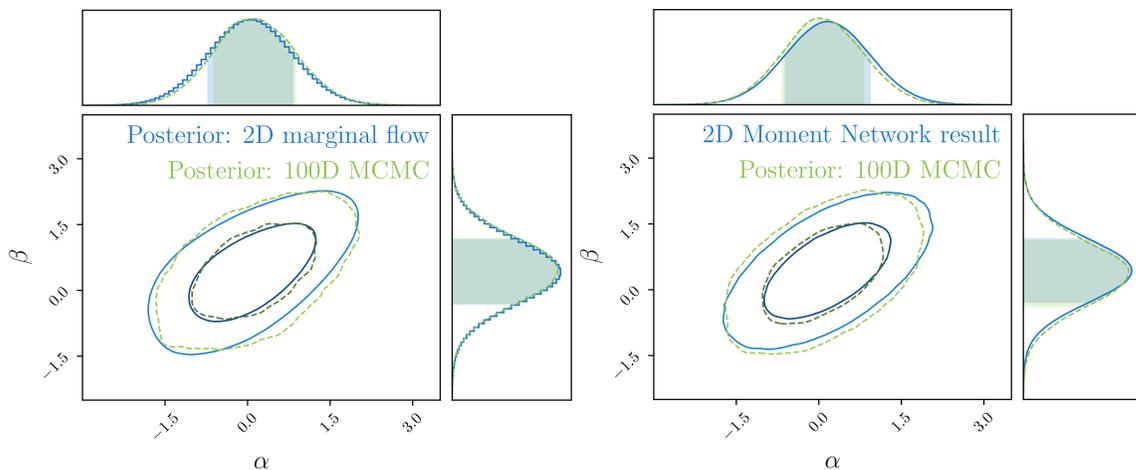
<sup>4</sup> *Center for Computational Astrophysics, Flatiron Institute, New York*

ArXiv:2011.05991 [github.com/NiallJeffrey/MomentNetworks](https://github.com/NiallJeffrey/MomentNetworks) niall.jeffrey [at] phys.ens.fr



## Highlights

- High-dimensional probability density estimation suffers from “*curse of dimensionality*”
- Propose direct estimation of lower-dimensional marginal posterior distributions
- **Marginal flow**: estimate joint-marginals of subsets of parameters
- **Moment Networks**: hierarchy of fast regression models compute increasing moments of lower-dimensional marginal posterior density
- **Beyond likelihood-free inference**: can also efficiently solve known MCMC problems
- **Demo**: high-dimensional inference: a) MCMC reference, b) Gravitational wave data model



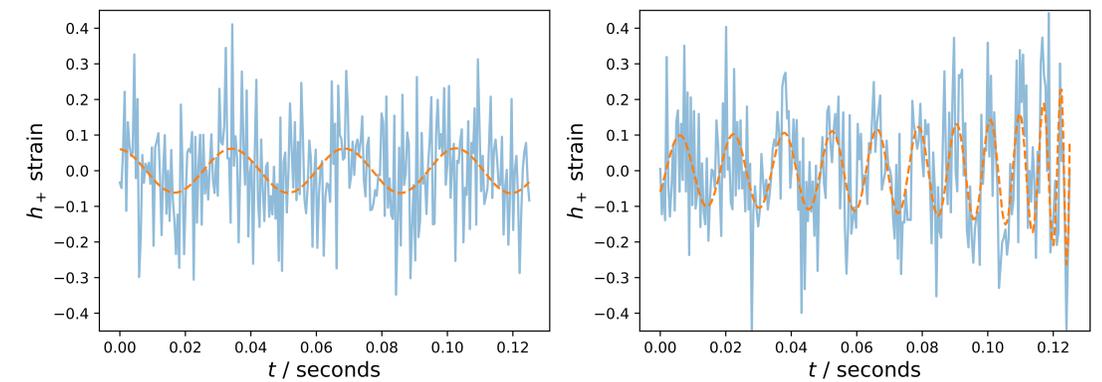
(A) 100-dimensional data model with known reference distribution evaluated with  $10^7$  MCMC samples. Direct 2D marginal posterior estimation using a masked autoregressive flow ensemble (left panel) and representation of 2D Moment Network result (right panel) both trained with  $8 \times 10^4$  simulations.

## Marginal flows

- Full  $p(x|\theta)$  density estimation for data  $x$  and parameters  $\theta$  intractable in high-dimension
- Make marginal densities the target of inference
- Estimate marginal posterior probability density for pairs of parameters  $\alpha, \beta \in \theta$  by minimizing  $-\sum_i \log q(\alpha_i, \beta_i | x_i)$
- Result  $q$  is an estimate (e.g. with normalizing flow) of marginal posterior  $p(\alpha, \beta | x)$  if training parameters  $\theta_i$  drawn from prior  $p(\theta)$  and used to simulate  $x_i$

**Markov chain Monte Carlo (MCMC) validation:**  
see Figure A

(B) Example simulated gravitational wave time series signals for the strain “+” polarization  $h_+$  with realistic LIGO-like noise.

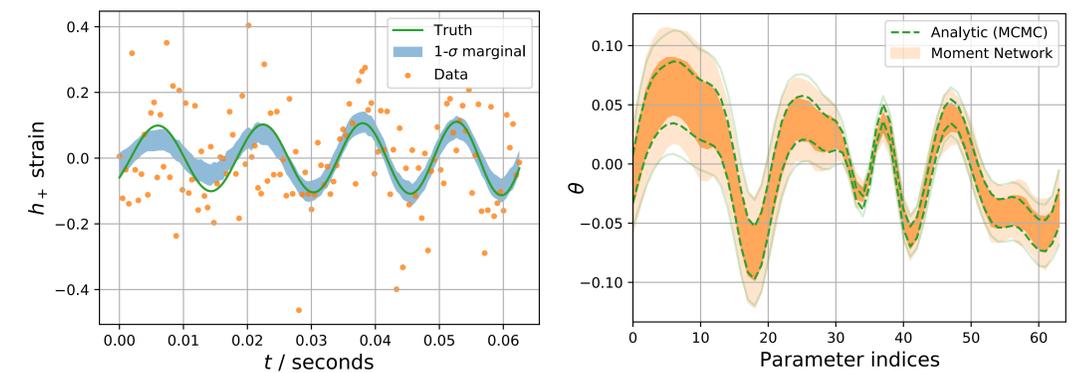


## LIGO-like gravitational wave time series

- Fig. B: two example simulated gravitational wave time series. The simplified signals (dashed orange) are  $\sim 0.12s$  intervals from the 1 second before a binary black hole merger
- Fig. C *left panel*: Moment Network estimate of the marginal mean and variance for each time step parameter
- Fig. C *right panel*: validation case of similar complexity, but with known likelihood
- Trained Moment Network accurately matches long-run MCMC chain, which validates our approach

(C) *Left panel*: Moment Network (MN) estimate of the 1- $\sigma$  per strain parameter.

*Right panel*: For each of 64 parameters (timestep), contours are marginal posterior  $\sigma$  from MN (shaded orange) and MCMC (dashed green).



## Cosmological applications

- **Mapping the Universe**: tractable simulation-based inference of high-dimensional cosmological fields & dark matter maps
- **Robustness**: Moment Networks use simpler architectures, reducing training failure risk and boosting inference speed
- **Cross-validation**: moments of estimated marginal posteriors should match those from Moment Networks
- **Evading MCMC**: Even when the likelihood can be sampled, marginal flows and Moment Networks still provide advantages, as many of the drawbacks of high-dimensional MCMC can be simply avoided